



NASA ExEP Resources Available for TDEM PIs

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Jet Propulsion Laboratory/California Institute of Technology

Pre-Proposal TDEM-17 Briefing Telecon

01/23/18

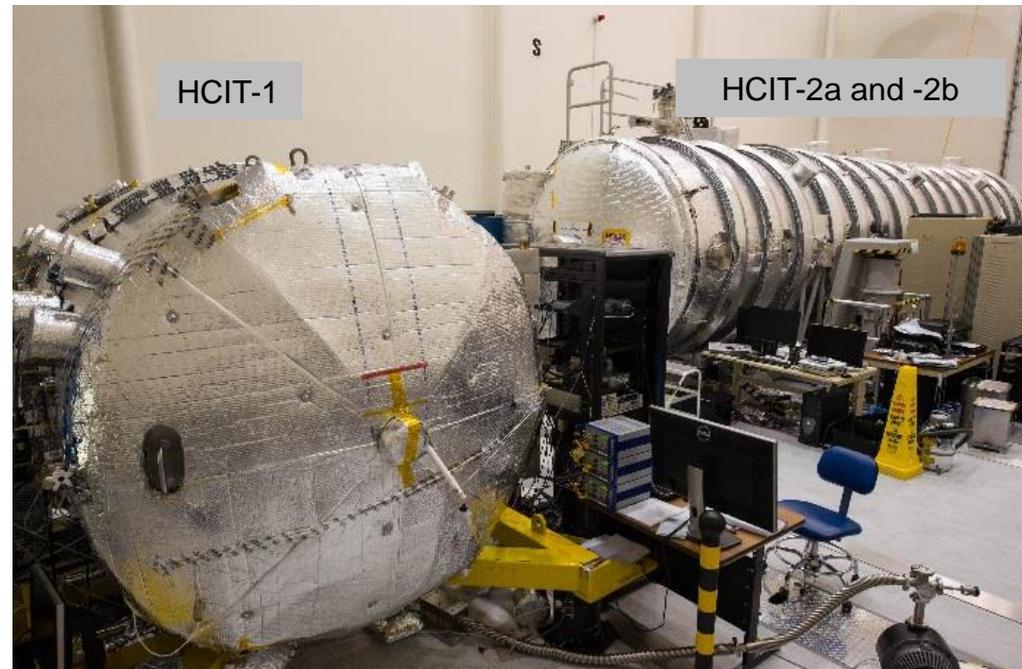
- This presentation provides an overview of the ExEP resources located at JPL available to support a TDEM-17 proposal.
- The available resources, if appropriate for your needs, may help you more efficiently meet your milestone goals and reduce your proposal costs and schedule.

Unavailable Resources

- HCIT-1 (dedicated to WFIRST)

Available Resources

- HCIT-2a (Decadal Survey Testbed)
 - Obscured, segmented apertures only
- HCIT-2b
- Vacuum Surface Gauge
- Microdevices Laboratory (MDL)





Gaining Access to the ExEP Resources at JPL



How to Request Use of ExEP Resources at JPL



Exoplanet Exploration Program

- **Submit preliminary Statement of Work (SOW) for use of ExEP resources to Brendan Crill no later than March 5, 2018.**
 - Follow SOW questionnaire on next page.
- **Schedule telecon with Brendan Crill between March 5 – 12, 2018 to discuss use of the resources of interest and to obtain costing guidelines.**
 - We will evaluate with the PI workforce, labor, and infrastructure access required across all received SOWs.
- **Brendan Crill will supply the proposal PI a Letter of Commitment for use of any ExEP resources.**
 - PIs are to include both the SOW and the Letter of Commitment in their proposal.
 - HCIT will provide workforce cost to set up testbeds; additional labor and unique procurements must be costed within the proposal.



SOW Questionnaire for Use of ExEP Resources



Exoplanet Exploration Program

- 1. Brief description of the proposed TDEM**
- 2. What resources are requested?**
- 3. Milestone (s) to be accomplished and performance goals**
- 4. Brief description of how the work will be conducted**
- 5. Period(s) and preferred dates, if any, over which the resource is requested, stating whether in vacuum or air for testbeds. Include any time required for preparatory work.**
- 6. A list of the personnel, expertise, and level of effort (if any) who will assist in the use of the resource.**
- 7. Any anticipated changes to the resource needed to accommodate your demonstrations.**
- 8. List of items needed for all testbed modifications. Identify items you will be procuring within your proposal's budget and provide approximate cost of needed items.**
 - a. Otherwise, state that no additional procurements will be necessary for the use of the infrastructure under consideration.
- 9. Provide any other relevant information or constraints.**



ExEP Technology Resources POC



Exoplanet Exploration Program

For questions concerning use of ExEP technology resources or requests for more detail contact:

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Jet Propulsion Laboratory

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Additional Slides



High Contrast Imaging Testbeds (HCITs)



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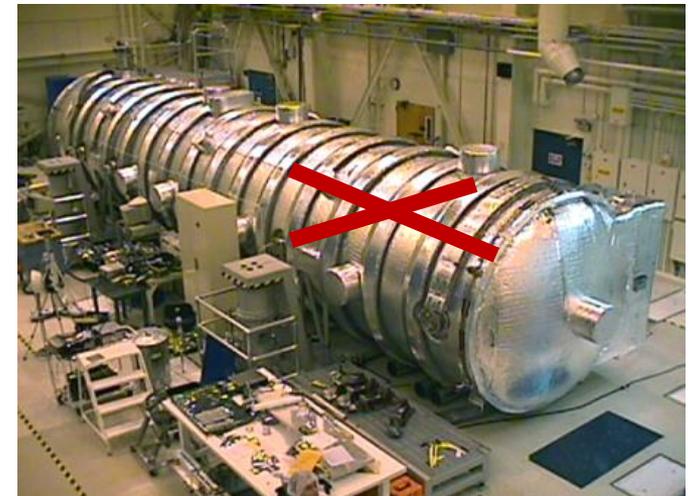
Test Facility

- Two vacuum chambers with 1 mTorr capability
- Seismically isolated, temperature-stabilized ~ 10 mK at RT.
- **Narrow or broad band coronagraph system demos**
 - Achieved 3×10^{-10} contrast (narrowband)
- **Fiber/Pinhole “Star” Illumination**
 - Monochromatic: 635, 785, 809, and 835 nm wavelengths
 - 2, 10, and 20% BW around 800 nm center
 - Medium and high power super-continuum sources
- **CCD camera ($5e^-$), 13 μm pixels**
- **Complete computer control with data acquisition and storage**
- **Coronagraph model validation & error budget sensitivities**
- **Remote access through FTP site**



HCIT-1

Single-testbed capacity (5'x8')

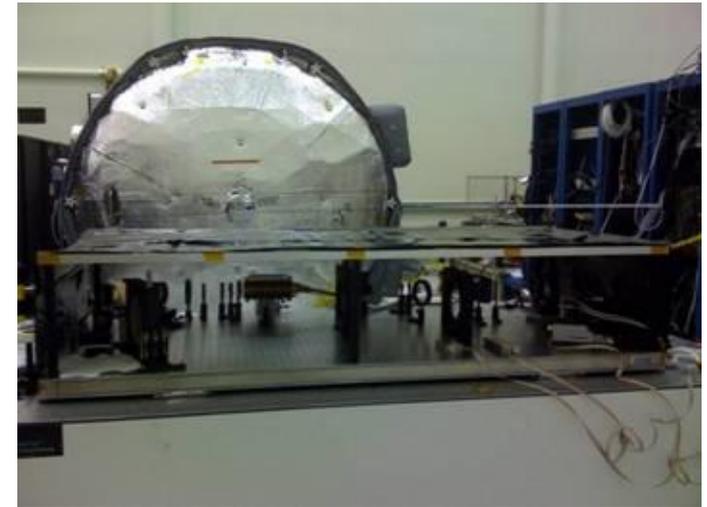


HCIT-2

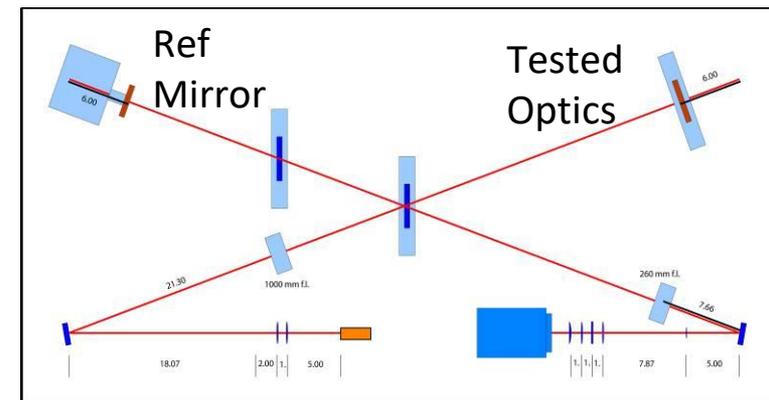
Two-testbed capacity (6'x10')

Purpose: Accurate surface error measurement and deformable mirror calibration.

- Demonstrated optical surface measurement accuracy: ≤ 100 pm rms
- Customized Michelson interferometer set-up
 - Reference mirror w/ absolute position feedback
 - Frequency stabilized laser source
- Dedicated algorithms for wavefront extraction over $> 10^6$ pixels



Vacuum Surface Gauge testbed



Surface Gauge optical layout

Purpose: Precision sub-micron materials fabrication and characterization

Advanced fabrication and characterization techniques

- Electron Beam Lithography
- Deep Reactive Ion Etching
- ICP Cryo Etching of Black Silicon microstructures
- Scanning Electron Microscopy
- Precision Optical Microscopy
- Atomic Force Microscopy
- 2D and 3D profilometry

Light suppression mask fabrication processes developed for:

- Micro dot patterned mask for JWST (Fig 1)
- Diffractive optical structures for spectrometer gratings and other computer generated holograms (Fig 2)
- Shaped pupil masks with fine structures and slits for transmission geometry (Fig 3)
- Shaped Pupil masks with black silicon structures in reflective aluminum background (Fig 4)
- LOWFS masks (Fig 5) incorporating a black silicon region (Fig 6) as well as shaped aperture through a silicon wafer
- Achromatic focal plane masks with deep diffractive structures (Fig 7)
- PIAACMC mask (Fig 8)
- Hybrid Lyot mask for WFIRST(Fig 9)
- Starshade mask for Princeton testbed (Fig 10)

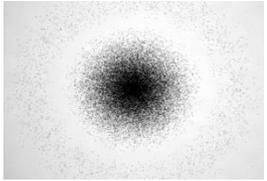


Figure 1. Microscope image (above) and AFM profile (below) of a micro dot patterned mask for JWST NIRCam coronagraph

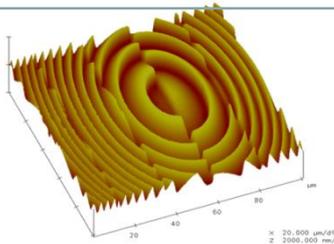


Figure 2. Diffractive optical devices

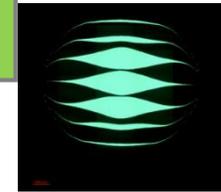
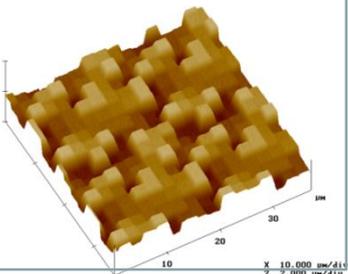


Figure 3. Transmissive slit SP mask

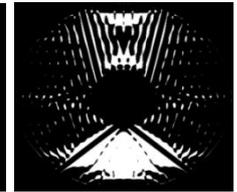


Figure 4. Reflective and absorptive SP mask

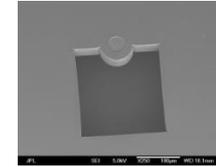


Figure 5. LOWFS mask

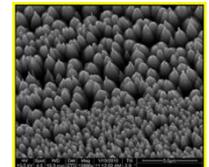


Figure 6. Black Si Microstructure

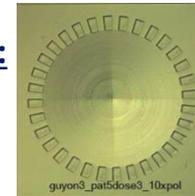


Figure 7 . Achromatic Focal Plane Masks (AFPM)

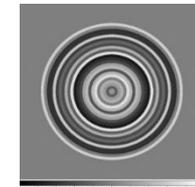
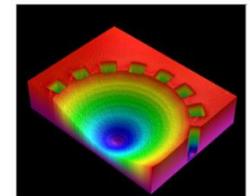


Figure 8 . PIAACMC mask

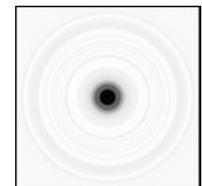


Figure 9 . Hybrid Lyot mask



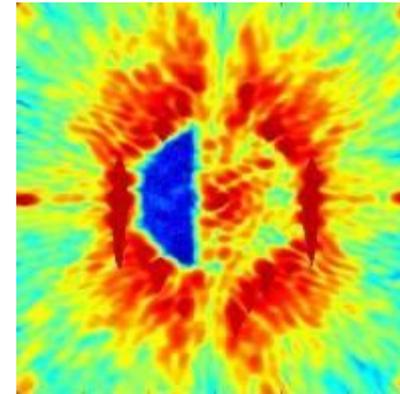
Figure 10 . Starshade mask

Nulling Algorithms

- Electric Field Conjugations (EFC) algorithms exist for single and dual DM control
- Demonstrated to $< 10^{-9}$ contrast and 20% bandwidth
- Coupled to HCIT coronagraph models and DM calibration data for optimal efficiency

Deformable Mirrors

- **Wavefront control and speckle nulling available with Xinetics PMN deformable mirrors.**
 - Format sizes: 32x32mm, 48x48, and 64x64 mm with 1 mm pitch and 500 nm stroke size.
 - Continuous fuse silica facesheet polished to $\lambda/100$ rms
 - Two-DM configurations available



Best Results to Date
Band-Limited Coronagraph :
6 e-10, @ 3 λ/D with 10% BW
2 e-9, @ 3 λ/D with 20% BW

Shaped-Pupil Coronagraph:
1.2 e-9, @ 4 λ/D with 2% BW
2.4 e-9, @ 4 λ/D with 10% BW

PIAA Coronagraph:
<1e-9, @ 2 λ/d with 0% BW

Vector Vortex Coronagraph:
<1e-9, @ 3 λ/d with 0% BW

EFC Nulling and current performance



Xinetics DM